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NONLINEAR EXCITATION OF SECOND SOUND IN QUANTUM SOLUTIONS DUE TO LIGHT WAVE ABSORPTION

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Nonlinear excitation of second sound in superfluid solutions $\text{He}^3 - \text{He}^4$ by light waves due to their absorption is discussed. Nonlinear wave equations that model the interaction of second-sound waves with light waves are derived, the expression for the nonlinear interaction length is obtained and an order-of-magnitude numerical estimate of the distance at which a second-sound wave could be amplified from a fluctuation level up to observable values is performed.

Nonlinear wave interactions that are essentially conditioned by wave absorption in a medium were first investigated in the frame of nonlinear optics (see Ref. [1] for a review). For the case of hydrodynamics this type of nonlinear processes was studied in Refs. [2, 3]. In the present paper the following absorption-induced nonlinear interaction is considered: two light waves E_1 and E_2 with slightly different frequencies propagate in a weak superfluid solution $\text{He}^3 - \text{He}^4$ at a small angle to each other and due to their absorption a second-sound wave with a frequency equal to the difference frequency of the light waves is excited and amplified.

The Hamiltonian approach will be used for describing second-sound wave propagation in quantum solutions. For superfluid solutions $\text{He}^3 - \text{He}^4$ Hamiltonian variables are three pairs of canonically conjugated variables (ρ, α) , (S, β) , (N, ξ) [4]. The meaning of these variables is as follows: ρ is the solution density; the quantity α determines the superfluid velocity $\mathbf{v}_s = \nabla \alpha$; S is the entropy density; N is the number of He^3 atoms per unit volume; the fluid unit volume momentum in the reference frame moving with the velocity \mathbf{v}_s is expressed via variables β and ξ as

$$\mathbf{j} = \rho_n (\mathbf{v}_n - \mathbf{v}_s) = S \nabla \beta + N \nabla \xi$$

where ρ_n, \mathbf{v}_n are the normal flow density and velocity. The mass flux density \mathbf{I} equals to

$$\mathbf{I} = \rho_s \mathbf{v}_s + \rho_n \mathbf{v}_n = \mathbf{j} + \rho \mathbf{v}_s,$$

ρ_s is the superfluid part density.

For investigating the nonlinear influence of optical waves on second sound-wave propagation we may limit ourselves with linear hydrodynamic equations [4] with a nonlinear optical source in the entropy equation

$$\begin{aligned} \frac{\partial \rho}{\partial t} &= -\Delta(S_0 \beta + N_0 \xi + \rho_0 \alpha); \\ \frac{\partial \alpha}{\partial t} &= -\mu; \quad \frac{\partial \beta}{\partial t} = -T; \quad \frac{\partial \xi}{\partial t} = -\zeta; \\ \frac{\partial S}{\partial t} &= -S_0 \Delta \left(\frac{S_0}{\rho_{n0}} \beta + \frac{N_0}{\rho_{n0}} \xi + \alpha \right) + \frac{Q}{T}; \\ \frac{\partial N}{\partial t} &= -N_0 \Delta \left(\frac{S_0}{\rho_{n0}} \beta + \frac{N_0}{\rho_{n0}} \xi + \alpha \right). \end{aligned} \quad (1)$$

In these equations μ is the chemical potential of the solution, ξ is the chemical potential of He^3 particles, T is the temperature, Q is the amount of heat emitted per unit volume and time unit due to light waves absorption. This quantity is proportional to the square modulus of the optical field $|E|^2$. The resonant for the second sound part of this quantity at the difference frequency is as follows

$$Q \sim \frac{c_l n \gamma_l}{8\pi} E_1 E_2^* e^{-i\Omega t},$$

where c_l is the light velocity, n – is the refraction index, γ_l is the light waves amplitude damping coefficient, $\Omega = \omega_1 - \omega_2$, where ω_1, ω_2 are the frequencies of the light waves E_1 and E_2 . Let the light waves with wave vectors \mathbf{k}_1 and \mathbf{k}_2 propagate at an angle \mathcal{G} to each other. In this case the wave vector \mathbf{q} of the excited second-sound wave is

$$|\mathbf{q}| \approx 2 \frac{\omega}{c} \sin \frac{\mathcal{G}}{2},$$

here $\omega \approx \omega_1 \approx \omega_2$.

We shall write Eqs. (1) in terms of canonical variables. For this it is convenient to introduce new variables ν, ψ, φ and η [4]: $\delta \mathbf{S} = \mathbf{S}_0 \nu$; $\psi = \mathbf{S}_0 \delta \mathbf{B}$; $\varphi = \rho_0 \delta \alpha$; $\delta \rho = \rho_0 \eta$. Here the following vector notations are used: $\mathbf{S} = (S, N)$, $\mathbf{B} = (\beta, \xi)$. The pairs (ν, ψ) и (η, φ) are also canonically conjugated variables. Using these variables we obtain Eqs. (1) in the form (we omit ‘0’ for undisturbed thermodynamic quantities):

$$\begin{aligned} \frac{\partial \eta}{\partial t} &= -\frac{1}{\rho} \Delta(\varphi + \psi); \\ \frac{\partial \nu}{\partial t} &= -\frac{1}{\rho} \Delta \left[\left(1 + \frac{\rho_s}{\rho_n} \right) \psi + \varphi \right] + \frac{1}{S + N} \frac{Q}{T}; \\ \frac{\partial \varphi}{\partial t} &= -\rho^2 \frac{\partial \mu}{\partial \rho} \eta - \rho \left(S \frac{\partial \mu}{\partial S} + N \frac{\partial \mu}{\partial N} \right) \nu; \\ \frac{\partial \psi}{\partial t} &= -\rho \left(S \frac{\partial T}{\partial \rho} + N \frac{\partial \xi}{\partial \rho} \right) \eta - \left[S \left(S \frac{\partial T}{\partial S} + N \frac{\partial \xi}{\partial S} \right) + N \left(S \frac{\partial T}{\partial N} + N \frac{\partial \xi}{\partial N} \right) \right] \nu. \end{aligned} \quad (2)$$

The Fourier-components of these variables for weak superfluid solutions $\text{He}^3 - \text{He}^4$ are expressed via second-sound normal coordinates b_q as

$$\begin{aligned} \eta_q &= \gamma B (b_q + b_{-q}^*), \\ \nu_q &= (\gamma - 1) B (b_q + b_{-q}^*), \\ iq \varphi_q &= \Gamma^{-1} \rho c_2 B (b_q - b_{-q}^*), \\ iq \psi_q &= -\Gamma^{-1} \rho c_2 B (b_q - b_{-q}^*), \\ \gamma &= -\frac{\sigma}{\rho} \left(\frac{\partial \rho}{\partial \sigma} \right)_{c,p} - \frac{c}{\rho} \left(\frac{\partial \rho}{\partial c} \right)_{\sigma,p}, \\ B &= \left(\frac{\rho_s}{\rho_n} \frac{\Omega}{2 \rho c_2^2} \right)^{1/2}, \end{aligned} \quad (3)$$

σ being the entropy per unit mass, c is the concentration, p is the pressure.

Eliminating from Eqs. (2) first φ , ψ , then the quantity η , taking into account the relations (3) and the fact that there is a nonlinear source in the equation for entropy, we come to the following equation for the slow variation of the second-sound amplitude v_q with distance:

$$\frac{dv_q}{dx} = \frac{\left[1 - c_2^{-2} \left(G + \rho \frac{\partial \mu}{\partial \rho}\right)\right] \frac{c_2 c_1 n}{T(S + N) 8\pi} E_1 E_2^*}{\frac{1}{c_2^2} \frac{\rho_s}{\rho_n} \left(DG - \frac{\partial \mu}{\partial \rho} F\right) + \left(1 + \frac{\rho_s}{\rho_n}\right) \left(-\chi G + \frac{F}{\rho}\right) - \gamma \rho \frac{\partial \mu}{\partial \rho} + D}. \quad (4)$$

In Eq. (4) the following notations are used:

$$\begin{aligned} D &= S \frac{\partial \mu}{\partial S} + N \frac{\partial \mu}{\partial N}; \\ G &= S \frac{\partial T}{\partial \rho} + N \frac{\partial \zeta}{\partial \rho}; \\ F &= S \left(S \frac{\partial T}{\partial S} + N \frac{\partial \zeta}{\partial S} \right) + N \left(S \frac{\partial T}{\partial N} + N \frac{\partial \zeta}{\partial N} \right). \end{aligned}$$

Let us make a numerical estimate of the distance l at which the second-sound intensity is amplified from fluctuation values up to observable values. From Eq. (4) we have

$$v_q(l) - v_q(0)_{fl} \approx v_q(l) = Al,$$

where A is the right-hand side of Eq. (4). For weak solutions the variable v_q is expressed via second-sound intensity as

$$v_q = \left(\frac{I}{2\rho c_2} \frac{\rho_s}{\rho_n} \right)^{1/2} \frac{1}{c_2},$$

and for the interaction length l we obtain

$$l = \left(\frac{I}{2\rho c_2} \frac{\rho_s}{\rho_n} \right)^{1/2} \frac{1}{c_2 A}.$$

It can be shown that the thermodynamic parameters entering the expression for A are of the following order of magnitude (taking into account in particular that the first-sound velocity c_1 is essentially higher than that of second sound)

$$\chi = \frac{c}{\rho} \frac{\partial \rho}{\partial c} \ll 1; \quad \gamma \sim \left(\frac{c_2}{c_1} \right)^2 \frac{\rho_n}{\rho_s} \ll 1 \text{ if } T \text{ is not close to } \lambda \text{ - point};$$

$$G + \rho \frac{\partial \mu}{\partial \rho} \sim \chi c_1^2; \quad DG \approx \left(N \frac{\partial \mu}{\partial N} \right)^2 \sim \chi^2 c_1^4; \quad F \frac{\partial \mu}{\partial \rho} \sim \frac{\rho_n}{\rho_s} c_1^2 c_2^2;$$

$$G \sim \chi c_1^2; \quad \frac{1}{\rho} F \sim \chi^2 c_1^2; \quad \rho \frac{\partial \mu}{\partial \rho} \sim \frac{\rho_n}{\rho_s} c_2^2; \quad D \sim \chi c_1^2.$$

Taking into account the above relations we can represent an order-of-magnitude value of the second-

sound excitation length as

$$l \sim \left(\frac{I}{2\rho c_2} \frac{\rho_s}{\rho_n} \right)^{1/2} \chi \frac{\rho_s}{\rho_n} \left(\frac{c_1}{c_2} \right)^2 \frac{T(S+N)}{\gamma_1 I_1},$$

where $I_1 \sim \frac{|E|^2}{8\pi} c_1 n$ is the light-wave intensity, $E \approx E_1 \approx E_2$.

To estimate numerically the distance l we will use experimental data given in Refs. [5, 6]:

$$T = 1.5^\circ \text{K}; \quad c_1 = 2.3 \cdot 10^4 \text{ cm/s}; \quad c_2 = 3 \cdot 10^3 \text{ cm/s}; \quad \frac{\rho_s}{\rho_n} \approx 3; \quad c < 0.1; \quad \chi \approx 0.02;$$

$$\rho = 0.14 \text{ g/cm}^3; \quad S = \rho\sigma \approx 5 \cdot 10^{21} \text{ cm}^{-3}; \quad N = \rho c / m_3 \approx 10^{21} \text{ cm}^{-3} \text{ (} m_3 \text{ is the mass of the He}^3 \text{ atom);}$$

$$\omega \approx 3 \cdot 10^{15} \text{ s}^{-1}; \quad \gamma_1 \approx (10^{-2} - 10^{-3}) \text{ cm}^{-1}.$$

For the light intensity $I_1 \sim 10^3 \text{ W/cm}^2$ the second sound can be amplified up to intensity $\sim 10^{-3} \text{ W/cm}^2$ at a distance $\sim 1 \text{ cm}$. Note, that from the point of view of an experiment in a superfluid orders-of-magnitude higher light-wave intensities are possible as well if using a pulse mode of light propagation (see, for ex. [7, 8]). In this case the excitation of second sound could be considerably more effective.

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